

Global Analysis of Nucleon Strange Form Factors at Low Q^2

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We perform a global analysis of all recent experimental data from elastic parity-violating electron scattering at low Q^2 . The values of the electric and magnetic strange form factors of the nucleon are determined at $Q^2 = 0.1 \text{ GeV}/c^2$ to be $G_E^s = -0.008 \pm 0.016$ and $G_M^s = 0.29 \pm 0.21$.

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The existence of a “sea” of quarks and antiquarks in the nucleon has been firmly established in deep-inelastic lepton scattering experiments as well as in the production of dilepton pairs (the Drell-Yan process). However, demonstrating the role of these $\bar{q}q$ pairs in the static electromagnetic properties of the nucleon has been a more elusive and difficult task.

As the lightest quark that contributes only to the $q\bar{q}$ sea, the strange quark provides a unique window on the role of the sea in the nucleon’s electromagnetic structure. As suggested by Kaplan and Manohar [1], knowledge of neutral current form factors, when combined with electromagnetic form factors, provides access to the contribution of strange quarks to these form factors. At low momentum transfers, the neutral current form factors can be determined through parity-violating (PV) electron scattering experiments [2, 3].

During the last decade, there has been dramatic progress in the study of the strange quark-antiquark contributions to the nucleon elastic electromagnetic form factors. A series of definitive PV electron scattering experiments along with several theoretical studies now provide a basis for extracting precision information on these strange quark contributions. In this work we report the results of a global analysis of all these experiments, including both the latest data obtained in experiments performed at the Jefferson Laboratory and appropriate theoretical input on radiative corrections, and obtain values for the strange electric and magnetic form factors of the nucleon at a four-momentum transfer $Q^2 = 0.1 \text{ GeV}/c^2$. We have also studied the sub-leading Q^2 dependence of these two form factors, and find that so far the data do not provide conclusive information.

I. STRANGE FORM FACTORS AND PARITY-VIOLATING ELECTRON SCATTERING

The nucleon vector strange form factors, G_E^s and G_M^s , characterize the contribution of the strange sea quarks to the nucleon electromagnetic form factors, and thereby

their contribution to the charge and magnetization distributions in the nucleon. With polarized electron facilities, G_E^s and G_M^s can be accessed by measuring the PV asymmetries in elastic e - p scattering, quasielastic e - d scattering, and elastic e - ^4He scattering [4]. In very general terms, the parity-violating asymmetry A_{PV} can be written as

$$A_{PV} = A_{nvs} + \eta_E G_E^s + \eta_M G_M^s, \quad (1)$$

where A_{nvs} is the “non-vector-strange” asymmetry (independent of G_E^s and G_M^s), and η_E and η_M are functions of kinematic quantities, nucleon electromagnetic form factors, and nuclear models (for non-hydrogen targets).

For elastic e - p scattering, the full form of the asymmetry is [4]

$$\begin{aligned} A_{PV}^p = & -\frac{G_F Q^2}{4\sqrt{2}\pi\alpha} \frac{1}{[\epsilon(G_E^p)^2 + \tau(G_M^p)^2]} \\ & \times \{(\epsilon(G_E^p)^2 + \tau(G_M^p)^2)(1 - 4\sin^2\theta_W)(1 + R_V^p) \\ & - (\epsilon G_E^p G_E^n + \tau G_M^p G_M^n)(1 + R_V^n) \\ & - (\epsilon G_E^p G_E^s + \tau G_M^p G_M^s)(1 + R_V^{(0)}) \\ & - \epsilon'(1 - 4\sin^2\theta_W)G_M^p G_A^e\}, \end{aligned} \quad (2)$$

with

$$\begin{aligned} \tau &= \frac{Q^2}{4M_p^2}, \quad \epsilon = \left(1 + 2(1 + \tau)\tan^2\frac{\theta}{2}\right)^{-1}, \\ \epsilon' &= \sqrt{\tau(1 + \tau)(1 - \epsilon^2)}, \end{aligned}$$

where M_p is the mass of the proton and θ is the electron scattering angle. In Eqn. 2, G_F and α are the Fermi and fine structure constants, respectively. Q^2 is the four momentum transfer. $G_{E,M}^{(p,n)}$ are the proton and neutron electric and magnetic form factors, while G_A^e is proton axial form factor seen by an electron. In order to extract contributions from $G_{E,M}^s$ to A_{PV}^p , one must include the effects of Standard Model (SM) $\mathcal{O}(\alpha)$ electroweak radiative corrections [5]. It is often useful to characterize these corrections in terms of ratios $R_{V,A}$ of the $\mathcal{O}(\alpha)$ hadronic vector (V) and axial vector (A) weak neutral current amplitudes to the corresponding tree-level amplitudes. The $R_V^p, R_V^n, R_V^{(0)}$ give these ratios for vector proton, neutron, and SU(3)-singlet amplitudes, respectively. In principle, their values can be obtained using the SM predictions

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for the effective electron-quark couplings C_{1q} given in [8]. However, the quoted C_{1q} do not include perturbative QCD contributions or coherent strong interaction effects in the radiative corrections associated with elastic scattering from a nucleon. A recent analysis of these effects has been given in Ref. [9] and up-dated in Ref. [10]. The latter work also gives an improved treatment of strong interaction contributions to the running of the weak mixing angle in the \overline{MS} renormalization scheme from its value at the Z -pole, $\hat{s}_Z^2 \equiv \sin^2 \hat{\theta}_W(M_Z)$, to the quantity appropriate for precise, low-energy neutral current experiments, $\sin^2 \hat{\theta}_W(0)$. All of these effects are included in the R_V given in Table I. The theoretical uncertainties in R_V^n and $R_V^{(0)}$ are less than one percent and have a negligible impact on our analysis, so we do not quote these errors in Table I. The theoretical error in $(1 - 4\sin^2 \theta_W)(1 + R_V^p)$ is ± 0.0008 [10], or slightly more than one percent. Since this error receives roughly equal contributions from the uncertainty in \hat{s}_Z^2 and from the $\mathcal{O}(\alpha)$ $Z\gamma$ box graph corrections, it is not appropriate to quote an uncertainty in R_V^p alone. The uncertainties associated with other strong interaction effects are sub-dominant. For the kinematic range of the PV experiments analyzed here, the R_V have a negligible impact on the Q^2 -dependence of A_{PV}^p and are taken to be constant. We use the conventional \overline{MS} renormalization scheme. Therefore, $\sin^2 \theta_W$ in Eqn. 2 and hereafter shall take its value as \hat{s}_Z^2 .

For reader's convenience, constant parameters (in some cases also the uncertainties) appearing in Eqn. 2 and later in Eqns. 3 and 4 are summarized in Table I.

Parameter	Value	Parameter	Value
α	1./137.03599911	g_A/g_V	-1.2695
\hat{s}_Z^2	0.23122(15)	$3F - D$	0.58(0.12)
G_F	$1.16637 \times 10^{-5}/\text{GeV}^2$	Δs	-0.07(0.06)
M_p	0.938272 GeV	R_V^p	-0.0520
Λ_A^2	1.00(0.04) (GeV/c) ²	R_V^n	-0.0123
		$R_V^{(0)}$	-0.0123

TABLE I: Summary of parameters (with some uncertainties in parentheses) in Eqns. 2, 3, and 4. The values of α , \hat{s}_Z^2 , G_F , M_p , g_A/g_V are taken directly from [8]. R_V^n , and $R_V^{(0)}$ are converted from C_{1q} parameters in [8]. R_V^p is derived from the proton weak charge given in Ref. [10]. We adopt the value and uncertainty of Λ_A^2 from [11], $3F - D$ from [12], and Δs from [13].

The effective axial form factor G_A^e receives a number of contributions and may be written as

$$G_A^e(Q^2) = G_D(Q^2) \times \left[\frac{g_A}{g_V} (1 + R_A^{T=1}) + \frac{3F - D}{2} R_A^{T=0} + \Delta s (1 + R_A^{(0)}) \right], \quad (3)$$

where

$$G_D(Q^2) = \frac{1}{(1 + Q^2/\Lambda_A^2)^2}, \quad (4)$$

parameterizes the Q^2 -dependence with a dipole form and Λ_A is the corresponding axial dipole mass. The ratio $-\frac{g_A}{g_V}$ is the isovector axial form factor of the nucleon at zero momentum transfer, which is precisely measured in the neutron beta decay. F and D are the octet baryon beta decay parameters, which can be determined by combining data from neutron and hyperon beta decays under the assumption of SU(3) flavor symmetry. Δs is the strange quark contribution to nucleon spin. Assuming a gentle evolution from the perturbative to the non-perturbative domain, this quantity can be obtained from inclusive, polarized deep inelastic lepton-nucleon scattering.

The ratios $R_A^{T=1}$, $R_A^{T=0}$, and $R_A^{(0)}$ characterize the effects of electroweak radiative corrections to the isovector, isoscalar, and SU(3) singlet hadronic axial vector amplitudes. Note that while $R_A^{T=1}$ and $R_A^{(0)}$ give the ratios of the $\mathcal{O}(\alpha)$ and tree-level hadronic axial vector neutral current amplitudes in the isovector and SU(3) singlet channels, respectively, $R_A^{T=0}$ does not have this interpretation since the tree-level isoscalar hadronic axial vector amplitude vanishes in the SM.

Conventionally, these quantities are divided into two pieces: the “one-quark” and “many-quark” contributions. The one-quark contributions correspond to renormalization of the effective vector electron-axial vector quark couplings, C_{2q} , and their values can be obtained from the SM predictions for these couplings given in Ref. [8]. The many-quark contributions include the so-called “anapole” effects as well as coherent strong interaction contributions to the radiative corrections. In contrast to the situation with the vector corrections, R_V , the *relative* importance of many-quark effects in the R_A can be quite pronounced. The small vector coupling of the electron to the Z -boson, $g_V^e = -1 + 4\sin^2 \hat{\theta}_W \sim -0.075$, leads to a suppression of the tree-level hadronic axial vector amplitude. However, g_V^e is absent from a variety of both one- and many-quark radiative corrections. Thus, one would expect the magnitudes of the R_A to be of order several percent, rather than the generic $\alpha/\pi \sim 0.3\%$ scale normally associated with electroweak radiative corrections. As a result, the impact of otherwise negligible strong interaction effects in the many-quark corrections, such as the anapole contributions, can be amplified [14].

An appropriate framework for treating the many-quark effects associated with physics at the hadronic scale is chiral perturbation theory. A comprehensive analysis of the anapole contributions to $R_A^{T=1}$ and $R_A^{T=0}$ has been carried out to chiral order p^3 in Ref. [15]. This analysis included both one-loop contributions associated with the octet of pseudoscalar mesons as well as the full set of low-energy constants (LECs) that arise at this chiral order. A generous theoretical range was assigned to the LECs, leading to a quoted theoretical uncertainty in the total R_A that is larger than the (logarithmically enhanced) one-quark corrections. The theoretical SM uncertainty is likely to be smaller.

The corresponding results, updated for the present value of the weak mixing angle, are given in Table I. The resulting prediction for G_A^e is consistent with both the results of the SAMPLE deuterium measurement [16, 17], which is particularly sensitive to the dominant isovector axial component, as well as other theoretical models for the anapole contributions [18, 19]. No evaluation of the “many-quark” contribution of $R_A^{(0)}$ has been made in the literature. We assume it is zero and assign the size of “one-quark” value for $R_A^{(0)}$ as its uncertainty.

	$R_A^{T=1}$	$R_A^{T=0}$	$R_A^{(0)}$
one-quark	-0.172	-0.253	-0.551
many-quark	-0.086(0.34)	0.014(0.19)	N/A
total	-0.258(0.34)	-0.239(0.20)	-0.55(0.55)

TABLE II: The “one-quark” [8] and “many-quark” [15] corrections to the axial charges, both in \overline{MS} , as well as the combined corrections.

The PV asymmetry for the neutron can be obtained by exchanging the “p” and “n” indices on nucleon form factors in Eqn. 2, and flipping the sign of the first isovector term in the expression for G_A^e in Eqn. 3. To first order, the PV asymmetry from a deuterium target is a cross-section weighted average of the proton and neutron asymmetries, which leads to an enhancement of the contribution of G_A^e and a suppression to the relative contribution due to G_E^s and G_M^s . Obviously a nuclear correction needs to be applied in the analysis. In this note, for the SAMPLE deuterium measurement, we shall adopt the asymmetry expression given in [16].

The ^4He nucleus is spin zero, parity even and isoscalar. The PV asymmetry takes a much simpler form [4]:

$$A_{PV}^{He} = \frac{G_F Q^2}{4\pi\sqrt{2}\alpha} \times \left(4\sin^2\theta_W(1 + R_V^{T=0}) + \frac{2(1 + R_V^{(0)})G_E^s}{G_E^p + G_E^n} \right), \quad (5)$$

where the isoscalar R_V factor is related to R_V^p and R_V^n as

$$R_V^{T=0} = \frac{R_V^n - (1 - 4\sin^2\theta_W)R_V^p}{4\sin^2\theta_W}. \quad (6)$$

II. EXPERIMENTAL DATA

In this section, the world data of PV elastic scattering within a Q^2 range from 0.07 to 0.5 (GeV/c) 2 will be summarized. These include SAMPLE-H [17, 20], SAMPLE-D [16, 17], HAPPEX-H-99 [21], HAPPEX-H-a [22], HAPPEX-He-a [23], HAPPEX-H-b and HAPPEX-He-b [24], PVA4-H-a [25], PVA4-H-b [26], and the first 14 Q^2 bins in G^0 forward angle [27]. The kinematics, targets, and the measured asymmetries in these experiments are summarized in Table III. In column A_{phys} ,

the first and second uncertainties for the G^0 data are the uncorrelated and correlated experimental uncertainties, respectively. The values of η_E and η_M are also listed in the table. In calculating them, we have adopted a recent parametrization of the nucleon electromagnetic form factors from Ref. [29]. For the SAMPLE deuterium measurement, the η_M is taken from Ref. [16], whereas its η_E is taken to be 1.79 according to the static approximation.

III. GLOBAL ANALYSIS

A. A_{nvs} and Theoretical Uncertainties

We shall now present a combined analysis of the world data aiming to extract G_E^s and G_M^s . A global fit, generally speaking, is obtained by simultaneously solving a set of equations

$$m_i \pm \sigma(m_i) = t_i(a_1, a_2, \dots) \pm \sigma(t_i), \quad (7)$$

where m_i and $t_i(a_1, a_2, \dots)$, respectively, are the measured and theoretical values for experiment i . In this expression, $\sigma(m_i)$ and $\sigma(t_i)$ are their uncertainties, and a_1, a_2, \dots are the free parameters one seeks to determine. In our case,

$$m_i = A_{phys}^i, t_i = A_{nvs} + \eta_E G_E^s + \eta_M G_M^s, \quad (8)$$

with G_E^s and G_M^s being the free parameters. In the previous section, we have discussed the value and uncertainty of A_{phys}^i , as well as η_E and η_M (Table III). For each measurement, the values of A_{nvs} can be also computed straightforwardly using the formalism in Sec. I. They are listed in Table IV. We again have used the parametrization of the nucleon electromagnetic form factors from Ref. [29]. As mentioned, the A_{nvs} for the SAMPLE deuterium measurement is calculated based on the asymmetry expression in [16] with the theoretical value of G_A^e .

The treatment of the theoretical uncertainties $\sigma(t_i)$ is more subtle. $\sigma(t_i)$ receives dominant contributions from the following sources: the nucleon axial form factor (G_A^e), nucleon electromagnetic form factors ($G_{E,M}^{(p,n)}$), and nuclear corrections. Theoretical uncertainties from a given source are correlated among different experiments. The uncertainty in G_A^e can be calculated based on the uncertainties in Tables I and I, and is dominated by the uncertainty of the “many-quark” electroweak radiative corrections on the R_A factors in Table I. For the nucleon electromagnetic form factors, based on the spread of the world data (see, e.g. Ref. [29]), we estimate their relative uncertainties as

$$\frac{\sigma(G_E^p)}{G_E^p} = 2.5\%, \frac{\sigma(G_M^p)}{G_M^p} = 1.5\%,$$

$$\frac{\sigma(G_E^n)}{G_E^n} = 15\% \text{ and } \frac{\sigma(G_M^n)}{G_M^n} = 1\%,$$

Experiment	Target	Q^2 (GeV/c) ²	θ_{lab} (deg)	A_{phys} (ppm)	η_E (ppm)	η_M (ppm)	Reference
SAMPLE-H	H ₂	0.1	144.4	-5.61 ± 1.11	2.07	3.47	[17, 20]
SAMPLE-D	D ₂	0.091	140.8	-7.77 ± 1.03	1.79	0.77	[16, 17]
HAPPEX-H-99	H ₂	0.477	12.3	-15.05 ± 1.13	56.89	22.62	[21]
HAPPEX-H-a	H ₂	0.099	6.0	-1.14 ± 0.25	9.55	0.76	[22]
HAPPEX-H-b	H ₂	0.109	6.0	-1.58 ± 0.13	10.59	0.93	[24]
HAPPEX-He-a	⁴ He	0.091	5.7	6.72 ± 0.87	20.19	0	[23]
HAPPEX-He-b	⁴ He	0.077	5.8	6.40 ± 0.26	16.56	0	[24]
PVA4-H-b	H ₂	0.108	35.52	-1.36 ± 0.32	10.08	1.05	[26]
PVA4-H-a	H ₂	0.23	35.45	-5.44 ± 0.60	22.56	5.07	[25]
G^0	H ₂	0.122	6.68	$-1.51 \pm 0.49 \pm 0.18$	11.96	1.17	[27]
G^0	H ₂	0.128	6.84	$-0.97 \pm 0.46 \pm 0.17$	12.60	1.30	[27]
G^0	H ₂	0.136	7.06	$-1.30 \pm 0.45 \pm 0.17$	13.46	1.48	[27]
G^0	H ₂	0.144	7.27	$-2.71 \pm 0.47 \pm 0.18$	14.32	1.66	[27]
G^0	H ₂	0.153	7.50	$-2.22 \pm 0.51 \pm 0.21$	15.31	1.89	[27]
G^0	H ₂	0.164	7.77	$-2.88 \pm 0.54 \pm 0.23$	16.53	2.19	[27]
G^0	H ₂	0.177	8.09	$-3.95 \pm 0.50 \pm 0.20$	17.99	2.58	[27]
G^0	H ₂	0.192	8.43	$-3.85 \pm 0.53 \pm 0.19$	19.69	3.07	[27]
G^0	H ₂	0.210	8.84	$-4.68 \pm 0.54 \pm 0.21$	21.77	3.71	[27]
G^0	H ₂	0.232	9.26	$-5.27 \pm 0.59 \pm 0.23$	24.37	4.60	[27]
G^0	H ₂	0.262	9.92	$-5.26 \pm 0.53 \pm 0.17$	28.00	5.99	[27]
G^0	H ₂	0.299	10.63	$-7.72 \pm 0.80 \pm 0.35$	32.60	7.99	[27]
G^0	H ₂	0.344	11.46	$-8.40 \pm 1.09 \pm 0.52$	38.40	10.89	[27]
G^0	H ₂	0.410	12.59	$-10.25 \pm 1.11 \pm 0.55$	47.28	16.10	[27]

TABLE III: A summary of the world data on PV elastic electron scattering within the range of $0.07 \text{ (GeV/c)}^2 < Q^2 < 0.5 \text{ (GeV/c)}^2$, including the average kinematics, targets, published asymmetries A_{phys} , as well as coefficients η_E and η_M . A_{phys} , η_E and η_M are in units of parts-per-million (ppm). The central kinematics for the two PVA4 measurements are obtained from [28]. For A_{phys} , the first and second uncertainties for the G^0 data are the uncorrelated and correlated experimental uncertainties, respectively.

respectively. This is consistent with the uncertainty assignment in Ref. [22], except that the uncertainty of G_E^n in [22] was assigned more conservatively to be 30%. Also note that we have made a simplifying assumption that these form factor uncertainties are “scaling” in nature, independent of the Q^2 . For an analysis with relatively small range of Q^2 , this is reasonable. Nuclear corrections are only relevant for non-hydrogen targets. For the two ⁴He measurements, according to Ref. [24], 3% is assigned as the fractional theoretical uncertainty of A_{nvs} . Nuclear corrections for the SAMPLE deuterium measurement have little impact on the final fit, and are therefore neglected. The theoretical uncertainties for the t_i in Eqn. 8 due to the different sources are summarized in the last six columns in Table IV. To be precise, the content of each column gives the change in $A_{nvs} + \eta_E G_E^s + \eta_M G_M^s$ when the source magnitude ($|G_A^e|$, $|G_{E,M}^{(p,n)}|$, or nuclear correction to ⁴He data) is increased by one standard deviation. Notice that the nucleon electromagnetic form factors also affect the value of η_E and η_M , therefore generate “pull terms” linear in G_E^s and G_M^s in Table IV. Such pull terms are neglected for the SAMPLE deuterium measurement.

B. Combined Analysis at $Q^2=0.1 \text{ (GeV/c)}^2$

As seen from Table III, a wealth of data exist with Q^2 in the vicinity of 0.1 (GeV/c)^2 , including SAMPLE-H, SAMPLE-D, HAPPEX-H-a, HAPPEX-H-b, HAPPEX-He-a, HAPPEX-He-b, PVA4-H-b, and low Q^2 data from G^0 . It is natural to first make a combined analysis at $Q^2 = 0.1$. To interpolate all data to a common Q^2 , we assume $G_E^s \propto Q^2$ and G_M^s is a constant [30]. That is, we replace $\eta_E G_E^s(Q^2) + \eta_M G_M^s(Q^2)$ by $\tilde{\eta}_E G_E^s(0.1) + \eta_M G_M^s(0.1)$, where $\tilde{\eta}_E = \eta_E \frac{Q^2}{0.1 \text{ (GeV/c)}^2}$. To simplify our notation, we shall use G_E^s and G_M^s hereafter to denote their values at $Q^2=0.1 \text{ (GeV/c)}^2$. In the (G_E^s, G_M^s) space, each measurement i provides a linear constraint as

$$\tilde{\eta}_E^i G_E^s + \eta_M^i G_M^s = A_{phys}^i - A_{nvs}^i. \quad (9)$$

In Fig. 1, each constraint is shown as a linear band in the (G_E^s, G_M^s) plane, where $\sigma(A_{phys}^i)$ (see Table III) and the theoretical uncertainty $\sigma(t_i)$ (see Table IV) have been combined in quadrature into an overall uncertainty. Somewhat arbitrarily, we include the 3 lowest Q^2 bins from G^0 data in this part of the analysis. For visual

Exp	Q^2 (GeV/c) ²	A_{nvs} ppm	Theoretical uncertainty $\sigma(t_i)$ (ppm)					
			G_A^e	G_E^p	G_M^p	G_E^n	G_M^n	N.C.
SAMPLE-H	0.1	-6.85	0.57	$0.06 + 0.03G_E^s - 0.03G_M^s$	$0.06 - 0.05G_E^s - 0.03G_M^s$	0.01	-0.05	0
SAMPLE-D	0.091	-8.37	0.60	0.05	0.02	0.02	-0.03	0
HAPPEX-H-99	0.477	-15.96	0.23	$0.38 + 0.07G_E^s - 0.52G_M^s$	$-0.03 - 0.90G_E^s - 0.02G_M^s$	0.49	-0.16	0
HAPPEX-H-a	0.099	-1.42	0.01	$0.04 - 0.15G_E^s - 0.03G_M^s$	$-0.01 - 0.05G_E^s + 0.01G_M^s$	0.05	-0.01	0
HAPPEX-H-b	0.109	-1.64	0.02	$0.05 - 0.16G_E^s - 0.04G_M^s$	$-0.01 - 0.06G_E^s + 0.01G_M^s$	0.06	-0.01	0
HAPPEX-He-a	0.091	7.50	0	$-0.47G_E^s$	0	$-0.12G_E^s$	0	0.23
HAPPEX-He-b	0.077	6.35	0	$-0.39G_E^s$	0	$-0.08G_E^s$	0	0.19
PVA4-H-b	0.108	-2.02	0.09	$0.06 - 0.14G_E^s - 0.04G_M^s$	$-0.02 - 0.07G_E^s + 0.01G_M^s$	0.05	-0.01	0
PVA4-H-a	0.23	-6.26	0.26	$0.17 - 0.13G_E^s - 0.15G_M^s$	$-0.03 - 0.26G_E^s + 0.02G_M^s$	0.18	-0.05	0
G^0	0.122	-1.94	0.02	$0.06 - 0.17G_E^s - 0.04G_M^s$	$-0.02 - 0.08G_E^s + 0.01G_M^s$	0.07	-0.02	0
G^0	0.128	-2.09	0.02	$0.06 - 0.17G_E^s - 0.05G_M^s$	$-0.02 - 0.09G_E^s + 0.01G_M^s$	0.07	-0.02	0
G^0	0.136	-2.29	0.02	$0.07 - 0.18G_E^s - 0.05G_M^s$	$-0.02 - 0.10G_E^s + 0.01G_M^s$	0.08	-0.02	0
G^0	0.144	-2.50	0.03	$0.07 - 0.18G_E^s - 0.06G_M^s$	$-0.02 - 0.11G_E^s + 0.01G_M^s$	0.09	-0.02	0
G^0	0.153	-2.74	0.03	$0.08 - 0.19G_E^s - 0.07G_M^s$	$-0.02 - 0.12G_E^s + 0.01G_M^s$	0.10	-0.02	0
G^0	0.164	-3.05	0.04	$0.09 - 0.19G_E^s - 0.08G_M^s$	$-0.02 - 0.13G_E^s + 0.01G_M^s$	0.11	-0.03	0
G^0	0.177	-3.43	0.04	$0.10 - 0.19G_E^s - 0.09G_M^s$	$-0.03 - 0.16G_E^s + 0.02G_M^s$	0.13	-0.03	0
G^0	0.192	-3.88	0.05	$0.11 - 0.19G_E^s - 0.10G_M^s$	$-0.03 - 0.18G_E^s + 0.02G_M^s$	0.14	-0.03	0
G^0	0.210	-4.46	0.06	$0.13 - 0.19G_E^s - 0.12G_M^s$	$-0.03 - 0.21G_E^s + 0.02G_M^s$	0.16	-0.04	0
G^0	0.232	-5.21	0.07	$0.15 - 0.19G_E^s - 0.15G_M^s$	$-0.04 - 0.25G_E^s + 0.02G_M^s$	0.19	-0.05	0
G^0	0.262	-6.30	0.08	$0.18 - 0.18G_E^s - 0.18G_M^s$	$-0.04 - 0.32G_E^s + 0.02G_M^s$	0.23	-0.06	0
G^0	0.299	-7.75	0.11	$0.21 - 0.15G_E^s - 0.23G_M^s$	$-0.04 - 0.40G_E^s + 0.02G_M^s$	0.27	-0.07	0
G^0	0.344	-9.65	0.14	$0.26 - 0.11G_E^s - 0.30G_M^s$	$-0.04 - 0.51G_E^s + 0.02G_M^s$	0.33	-0.09	0
G^0	0.410	-12.70	0.19	$0.32 - 0.03G_E^s - 0.40G_M^s$	$-0.04 - 0.69G_E^s + 0.00G_M^s$	0.41	-0.12	0

TABLE IV: A summary of A_{nvs} and the theoretical uncertainties of individual measurements. The uncertainties are grouped by six different sources: G_A^e , G_E^p , G_M^p , G_E^n , G_M^n and the nuclear corrections (N.C.) to ^4He data. See text for details.

clarity, they are combined into a single constraint as

$$A_{phys}^i - A_{nvs}^i = 0.84 \pm 0.34 = 16.38G_E^s + 1.32G_M^s, \quad (10)$$

which is shown as the solid brown band in Fig. 1. From the figure, one sees that the agreement among different measurements is generally good. The G^0 and PVA4 appear to be offset from the HAPPEX-H-b measurement, but they nevertheless agree within 2σ . As explained in Sec. I, the SAMPLE deuterium measurement (dashed red band) has much less sensitivity to G_E^s and G_M^s .

The 10 different measurements (3 from G^0 , and the other 7 from separate experiments) above provide redundancy in the joint determination of (G_E^s, G_M^s) . To solve for (G_E^s, G_M^s) and determine the confidence contours, we follow the standard least square procedure (see, e.g. Ref [31]). Specifically, we rearrange Eqn. 9 into the form of Eqn. 7 as

$$m_i = t_i(G_E^s, G_M^s) \pm \sigma_i + \sum_{k=0}^6 \pm \beta_{i,k}. \quad (11)$$

where m_i and t_i are given in Eqn. 8, and σ_i is the uncorrelated experimental uncertainty. $\beta_{i,k}$ denotes the correlated uncertainty for measurement i with “source index” k . In our case, $\beta_{i,0}$ equals the correlated experimental uncertainty for the G^0 data and 0 for other experiments,

and $\beta_{i,k=1,2,3,4,5,6}$ are the correlated theoretical uncertainties for each measurement i due to different sources (Table IV). Then for each given pair of (G_E^s, G_M^s) , the χ^2 is calculated as

$$\chi^2 = \sum_i \sum_j (m_i - t_i)(V^{-1})_{ij}(m_j - t_j), \quad (12)$$

where i and j are indices of the measurements, and V is the variance matrix with $V_{ij} = \sigma_i^2 \delta_{ij} + \sum_{k=0}^6 \beta_{i,k} \beta_{j,k}$.

It has been shown in Ref. [31] that the χ^2 constructed this way satisfies the standard χ^2 distribution, and the solution (best fit) can be found by minimizing this χ^2 . Applying this technique to the 10 measurements in Fig. 1, we obtain

$$G_E^s = -0.006 \pm 0.016, \quad G_M^s = 0.33 \pm 0.21 \quad (13)$$

with a correlation coefficient of -0.83 between the two, and a minimum $\chi_{min}^2 = 9.90$ with 8 degrees of freedom. Note that the uncertainties above are 1σ ($\Delta\chi^2 = 1$) “marginalized” uncertainties corresponding to the projections of the error ellipse onto the two axes. That is, for a given value of G_M^s , the χ^2 is minimized by varying G_E^s and vice versa. The range defined by this uncertainty for a given parameter corresponds precisely to

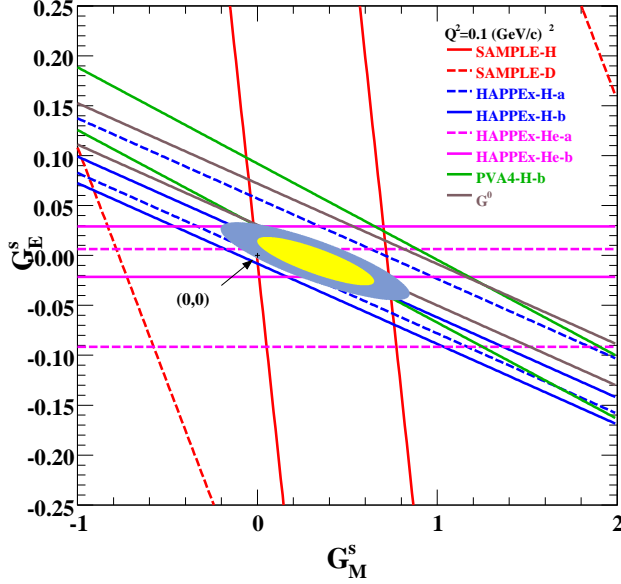


FIG. 1: (Color online) The world data constraints on (G_E^s, G_M^s) at $Q^2 = 0.1 \text{ (GeV/c)}^2$. The form factors of Kelly are used. Different bands in the plot represent: SAMPLE-H [20] (solid red), SAMPLE-D [16] (dashed red), HAPPEX-H-a [22] (dashed blue), HAPPEX-H-b [24] (solid blue), HAPPEX-He-a [23] (dashed pink), HAPPEX-He-b [24] (solid pink), PVA4-H-b [26] (solid green), and the lowest three Q^2 bins in G^0 forward angle [27] (solid brown). The yellow and gray blue (dark) ellipses represents 68.27% ($\Delta\chi^2 = 2.3$) and 95% ($\Delta\chi^2 = 5.99$) confidence contours around the point of maximum likelihood at $(G_E^s = -0.006, G_M^s = 0.33)$. The black cross represents $G_E^s = G_M^s = 0$.

68.27% of confidence interval of that parameter [32]. On the other hand, for the two parameters $(G_E^s \text{ and } G_M^s)$ that are jointly determined, the 68.27% confidence region is instead defined by $\Delta\chi^2 = 2.3$ contour [8]. To demonstrate the precision of the fit, we plot the 68.27% ($\Delta\chi^2 = 2.30$) and 95% ($\Delta\chi^2 = 5.99$) joint confidence levels in Fig. 1 as the yellow (light) and gray (dark) ellipses, respectively. $(G_E^s, G_M^s) = (0, 0)$ yields a $\Delta\chi^2$ of 4.81 ($\int_{4.8}^{\infty} P_{\chi^2} d\chi^2 = 9.0\%$ where P_{χ^2} is the χ^2 probability distribution function [32]), which is depicted as the black cross in the figure.

C. Inclusion of higher Q^2 data from G^0 , PVA4, and HAPPEX

The analysis presented so far has focused on the vicinity of $Q^2=0.1 \text{ (GeV/c)}^2$. It is desirable to extend the analysis by including data at higher Q^2 from G^0 , PVA4, and HAPPEX. As a first attempt, the previous assumption that $G_E^s \propto Q^2$ and G_M^s is a constant was adopted. We shall refer to such a fit as the first order fit. Using the same χ^2 construction as in Eqns. 11 and 12, we started by fitting the data in Table III with $Q^2 < 0.11 \text{ (GeV/c)}^2$ (up

to PVA4-H-b), and then systematically included more data in the fit with increasing Q^2 . In Table V, the resulting (G_E^s, G_M^s) and the fit quality are summarized. Compared to the results in Sec. IIIB, extending the fit

Q^2 (GeV/c) ²	Experiment	G_E^s	G_M^s	χ^2_{min}/ν	Prob (%)
<0.11 (GeV/c) ²		-0.008(0.016)	0.26(0.22)	4.7/5	45.5
0.122	G^0	-0.007(0.016)	0.27(0.22)	4.9/6	55.4
0.128	G^0	-0.006(0.016)	0.30(0.22)	8.4/7	29.7
0.136	G^0	-0.006(0.016)	0.33(0.21)	9.9/8	27.2
0.144	G^0	-0.007(0.016)	0.30(0.21)	13.0/9	16.1
0.153	G^0	-0.007(0.016)	0.30(0.21)	13.0/10	22.1
0.164	G^0	-0.008(0.016)	0.29(0.21)	13.8/11	24.4
0.177	G^0	-0.010(0.016)	0.24(0.21)	19.5/12	7.7
0.192	G^0	-0.010(0.016)	0.22(0.21)	20.3/13	8.9
0.210	G^0	-0.012(0.016)	0.19(0.21)	22.1/14	7.6
0.230	PVA4	-0.013(0.016)	0.22(0.20)	22.3/15	10.1
0.232	G^0	-0.014(0.016)	0.21(0.20)	23.1/16	11.0
0.262	G^0	-0.012(0.016)	0.24(0.19)	24.7/17	10.2
0.299	G^0	-0.014(0.016)	0.23(0.19)	26.0/18	10.0
0.344	G^0	-0.014(0.016)	0.24(0.19)	26.1/19	12.7
0.410	G^0	-0.013(0.015)	0.27(0.19)	27.1/20	13.1
0.477	HAPPEX-H-99	-0.015(0.015)	0.25(0.19)	28.6/21	12.4

TABLE V: Results of the first order fit ($G_E^s \propto Q^2, G_M^s = \text{constant}$). Rows are ordered by the Q^2 of the data. Each row gives the results of the fit that includes all data in Table III up to, and including, the given Q^2 . Row “<0.11 (GeV/c)²” represents the fit with low Q^2 data included up to PVA4-H-b. Columns “ χ^2_{min}/ν ” and “Prob” are the reduced χ^2 and the probability ($\equiv \int_{\chi^2_{min}}^{\infty} P_{\chi^2} d\chi^2$) of the fit.

to higher Q^2 gives consistent values for G_E^s and G_M^s . One also notices that for Q^2 beyond 0.164 (GeV/c)^2 , the fit quality (χ^2_{min}/ν and χ^2 probability) deteriorates significantly, implying that our lowest order model is no longer able to capture the true Q^2 variation in these two quantities.

To better characterize the data at higher Q^2 , therefore, one needs to introduce higher order Q^2 dependence to G_E^s and/or G_M^s . For $\sqrt{Q^2}$ sufficiently below the mass of the kaon, chiral perturbation theory provides a systematic framework for characterizing the Q^2 -dependence of nucleon form factors. In this context the strange magnetic moment, $\mu_s \equiv G_M^s(0)$, arises at chiral order p^2 . The sub-leading Q^2 -dependence of $G_M^s(Q^2)$ – sometimes called the strange magnetic radius – is nominally $\mathcal{O}(p^4)$ and is determined by a combination of chiral loop contributions and a corresponding LEC. Loop effects also generate an order p^3 contribution to the strange magnetic radius that is free of any new parameters [33]. However, this contribution is substantially canceled by the $\mathcal{O}(p^4)$ loop corrections, resulting in a strong dependence on the $\mathcal{O}(p^4)$ LEC [35]. In contrast, G_E^s starts off at chiral order p^3 , while the sub-leading Q^2 -dependence (and corresponding LEC) arises at $\mathcal{O}(p^5)$.

Based on these considerations, we extend the previous

analysis to include all known, sub-leading Q^2 -dependence in $G_{M,E}^s$ through $\mathcal{O}(p^4)$. In practice, doing so amounts to including one new constant in our fit associated with the strange magnetic radius. Since we are interested in the implications for the values of $G_{M,E}^s$ at $Q^2 = 0.1$ (GeV/c) 2 , we expand about this value of Q^2 rather than about $Q^2 = 0$, *viz*,

$$G_E^s(Q^2) = G_E^s \frac{Q^2}{0.1}, \quad G_M^s(Q^2) = G_M^s + \mu'_s(Q^2 - 0.1), \quad (14)$$

where G_E^s and G_M^s again are used to represent their values at $Q^2 = 0.1$ (GeV/c) 2 . This will be referred to as the second order fit. As in the previous section, each row in Table III now leads to a new constraint in the form of Eqn. 11, with G_E^s and G_M^s parametrized by Eqn. 14. Again using Eqns. 11 and 12, we constructed a proper χ^2 function and performed the least square fit for (G_E^s, G_M^s, μ'_s) by including higher Q^2 data.

The results of this procedure are summarized in Table VI. Several observations can be made from the ta-

Q^2 (GeV/c) 2	G_E^s	G_M^s	μ'_s (GeV/c) $^{-2}$	χ^2_{min}/ν	Prob (%)
<0.11	-0.002(0.017)	0.37(0.25)	-23.7(27.5)	3.9/4	41.3
0.122	-0.006(0.017)	0.30(0.24)	-5.9(19.9)	4.8/5	43.6
0.128	-0.011(0.017)	0.20(0.23)	15.7(13.9)	7.1/6	30.7
0.136	-0.011(0.016)	0.21(0.23)	15.7(9.5)	7.1/7	41.4
0.144	-0.007(0.016)	0.30(0.22)	0.1(6.9)	13.0/8	11.0
0.153	-0.007(0.016)	0.30(0.22)	-0.1(5.3)	13.0/9	16.1
0.164	-0.006(0.016)	0.32(0.22)	-2.3(4.0)	13.5/10	19.9
0.177	-0.005(0.016)	0.35(0.22)	-5.8(2.7)	15.0/11	18.4
0.192	-0.006(0.016)	0.33(0.22)	-4.2(2.0)	15.7/12	20.6
0.210	-0.006(0.016)	0.33(0.22)	-3.6(1.4)	15.9/13	25.3
0.230	-0.012(0.016)	0.35(0.21)	-1.5(1.0)	20.1/14	12.6
0.232	-0.013(0.016)	0.35(0.21)	-1.6(0.9)	20.2/15	16.5
0.262	-0.012(0.016)	0.29(0.21)	-0.4(0.7)	24.4/16	8.2
0.299	-0.012(0.016)	0.31(0.21)	-0.6(0.6)	24.8/17	10.0
0.344	-0.012(0.016)	0.28(0.20)	-0.3(0.5)	25.7/18	10.8
0.410	-0.014(0.016)	0.25(0.20)	0.1(0.3)	27.0/19	10.4
0.477	-0.014(0.015)	0.28(0.20)	-0.1(0.2)	28.4/20	10.1
<0.11 and 0.477	-0.008(0.016)	0.26(0.22)	-0.3(0.3)	4.7/5	45.7

TABLE VI: Results of the second order fit ($G_E^s \propto Q^2$ and G_M^s linear in Q^2). See the caption of Table V for a description of the table content. Row “< 0.11” represents the fit using low Q^2 data up to PVA4-H-b. The data at $Q^2=0.230$ and 0.477 (GeV/c) 2 are from PVA4 and HAPPEX-H-99, respectively, and all others are from G^0 . Uncertainties of G_E^s , G_M^s , and μ'_s are “marginalized” 1σ uncertainties. The last row represents the fit by using only the low Q^2 (<0.11 (GeV/c) 2) and HAPPEX-99 data.

ble. First, compared to the first order fits, the additional free parameter μ'_s does not improve the fit quality significantly. Reasonable fits can be obtained up to Q^2 of 0.210 (GeV/c) 2 , beyond which the “flexibility” of our fit model again seems to be inadequate to describe the data. Second, the best values of G_E^s and G_M^s are very similar to those obtained from the first order fit (Table V).

Third, the uncertainties of G_E^s and G_M^s , as compared to those from the first order fit, are slightly larger due to the additional parameter μ'_s . Fourth, if we ignore the fit quality, and simply examine the mean and uncertainty of μ'_s , it is large and uncertain until the fit range goes up to 0.262 (GeV/c) 2 . This is expected, since it is difficult to determine the slope parameter with insufficient “lever arm”. Also, fits beyond 0.262 (GeV/c) 2 suggests a gentle μ'_s . (To illustrate this point, in the last row in Table VI we show the fit results by using only the low Q^2 (<0.11 (GeV/c) 2) and HAPPEX-99 data, which yield a consistent picture as described above.) However, the fit quality for $Q^2 > 0.210$ (GeV/c) 2 prevents us from making a strong statement about μ'_s here.

For completeness, we also investigated the impact of including a Q^4 term in G_E^s (corresponding to chiral order p^5). The resulting fit quality does not improve substantially with the inclusion of data with Q^2 beyond ~ 0.2 (GeV/c) 2 , and both second order parameters in the form factors are poorly constrained. Nevertheless, the resulting $G_{E,M}^s$ are consistent with those obtained above.

Based on these considerations, we choose to use the first order fit up to $Q^2 = 0.164$ (Table V) as our final results in this analysis:

$$G_E^s = -0.008 \pm 0.016, \quad G_M^s = 0.29 \pm 0.21, \quad (15)$$

with a correlation coefficient of -0.85 between the two, and $\chi^2_{min}/\nu = 13.8/11$. Again, the uncertainties of the two form factors are marginalized 1σ uncertainties corresponding to the projections of $\Delta\chi^2 = 1$ contour. This result is in good agreement with Eqn. 13, for which only the lowest three Q^2 bins of G^0 data were included. For G_M^s alone, the one-side confidence integral for a negative G_M^s ($[-\infty, 0]$) is 12.3%.

IV. CONCLUSION

A combined analysis of the world PV electron scattering data has been performed to extract the nucleon strange electric and magnetic form factors G_E^s and G_M^s at $Q^2 = 0.1$ (GeV/c) 2 . Our treatment is similar to that of Ref. [36], but utilizes all available low Q^2 data including the recent HAPPEX results [22, 23] and incorporates the uncertainties in the electromagnetic and axial form factors. We find that the agreement among different measurements is good and we obtain fits with acceptable χ^2 . Using a simple parametrization of the Q^2 variation in G_E^s and G_M^s , a satisfactory global fit can be obtained up to a $Q^2 \sim 0.2$ (GeV/c) 2 . At $Q^2 = 0.1$ (GeV/c) 2 , the confidence integral for $G_M^s < 0$ is 12.3%, so substantially negative values of G_M^s are highly disfavored by the fit. In addition, our best fit is consistent with $G_E^s = 0$ at $Q^2=0.1$ (GeV/c) 2 .

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- [1] D. Kaplan and A. Manohar, Nucl. Phys. B **310**, 527 (1988).
 - [2] R. D. McKeown, Phys. Lett. B **219**, 140 (1989).
 - [3] D. H. Beck, Phys. Rev. D **39**, 3248 (1989).
 - [4] M. J. Musolf *et al.*, Phys. Rep. **239**, No. 1 (1994).
 - [5] For recent analyses of isospin-breaking corrections, see Refs. [6, 7].
 - [6] B. Kubis and R. Lewis, Phys. Rev. C **74**, 015204 (2006) [arXiv:nucl-th/0605006].
 - [7] M. Viviani *et al.*, arXiv:nucl-th/0703051.
 - [8] W. M. Yao *et al.*, Journal of Physics G **33**, 1 (2006).
 - [9] J. Erler, A. Kurylov and M. J. Ramsey-Musolf, Phys. Rev. D **68**, 016006 (2003) [arXiv:hep-ph/0302149].
 - [10] J. Erler and M. J. Ramsey-Musolf, Phys. Rev. D **72**, 073003 (2005) [arXiv:hep-ph/0409169].
 - [11] H. Budd, A. Bodek and J. Arrington, arXiv:hep-ex/0308005.
 - [12] B. W. Filippone and Xiang-Dong Ji, Adv. Nucl. Phys. **26**, 1 (2001).
 - [13] D. Adams *et al.* [Spin Muon Collaboration (SMC)], Phys. Rev. D **56**, 5330 (1997).
 - [14] The one-quark contributions can be further enhanced by large logarithms associated with running between the weak and hadronic scales, leading to relative corrections of order ten percent or more.
 - [15] S. L. Zhu, S. J. Puglia, B. R. Holstein and M. J. Ramsey-Musolf, Phys. Rev. D **62**, 033008 (2000).
 - [16] T. M. Ito *et al.*, Phys. Rev. Lett. **92**, 102003 (2004).
 - [17] E. J. Beise, M. L. Pitt and D. T. Spayde, Prog. Part. Nucl. Phys. **54**, 289 (2005).
 - [18] D. O. Riska, Nucl. Phys. A **678**, 79 (2000) [arXiv:hep-ph/0003132].
 - [19] C. M. Maekawa and U. van Kolck, Phys. Lett. B **478**, 73 (2000) [arXiv:hep-ph/0006161].
 - [20] D. T. Spayde *et al.* (SAMPLE), Phys. Rev. Lett. **84**, 1106 (2000).
 - [21] K. A. Aniol *et al.* (HAPPEX), Phys. Rev. C **69**, 065501 (2004).
 - [22] K. A. Aniol *et al.* (HAPPEX), Phys. Lett. B **635**, 275 (2006).
 - [23] K. A. Aniol *et al.* (HAPPEX), Phys. Rev. Lett. **96**, 022003 (2006).
 - [24] A. Acha *et al.* (HAPPEX), Phys. Rev. Lett. **98**, 032301 (2007).
 - [25] F. E. Maas *et al.* (PVA4), Phys. Rev. Lett. **93**, 022002 (2004).
 - [26] F. E. Maas *et al.* (PVA4), Phys. Rev. Lett. **94**, 152001 (2005).
 - [27] D. S. Armstrong *et al.* (G^0), Phys. Rev. Lett. **95**, 092001 (2005).
 - [28] J. Arvieux, F. E. Maas, private communications.
 - [29] J. J. Kelly, Phys. Rev. C **70**, 068202 (2004).
 - [30] $G_E^s(0) \equiv 0$, since there is no net strangeness in nucleon.
 - [31] D. Stump *et al.*, Phys. Rev. D **65**, 014012 (2001).
 - [32] P. R. Bevington and D. K. Robinson, *Data Reduction and Error Analysis for the Physical Sciences*, second ed., McGraw-Hill, Boston, 1992.
 - [33] T. R. Hemmert, U. G. Meissner and S. Steininger, Phys. Lett. B **437**, 184 (1998) [arXiv:hep-ph/9806226].
 - [34] T. R. Hemmert, B. Kubis and U. G. Meissner, Phys. Rev. C **60**, 045501 (1999) [arXiv:nucl-th/9904076].
 - [35] H. W. Hammer, S. J. Puglia, M. J. Ramsey-Musolf and S. L. Zhu, Phys. Lett. B **562**, 208 (2003) [arXiv:hep-ph/0206301].
 - [36] R. D. Young, J. Roche, R. D. Carlini, A. W. Thomas, Phys. Rev. Lett. **97**, 102002 (2006).